Explicit Symmetries of the Kepler Problem

Horst Knörrer, ETH Zürich

The motion of a point mass q(t) in the gravitation field of a mass that is fixed at the origin is described by the Kepler Hamiltonian

$$H(q,p) = \frac{1}{2} ||p||^2 - \frac{\mu}{||q||}$$

Here, μ is determined by the gravitational constant and the two masses. Given an energy E, one regularizes the motion on the energy level hypersurface $\{(q,p) \mid H(q,p) = E\}$ by introducing the "eccentric anomaly" as new independent variable. If $E \neq 0$, the rescaled problem is the Hamiltonian flow of the Hamiltonian $\tilde{H}(q,p) = \frac{1}{\sqrt{2|E|}} ||q|| (||p||^2 - 2E) - \frac{2\mu}{\sqrt{2|E|}}$ on the hypersurface $\{(q,p) \mid \tilde{H}(q,p) = 0\}$. See [CB], II.5.10b. Scaling q by the factor $\sqrt{2|E|}$ and p by the factor $\frac{1}{2|E|}$ allows to reduce the discussion to the case 2|E| = 1. That is, we consider the Hamiltonian flows of

$$K_{\pm}(q,p) = ||q|| (||p||^2 \pm 1) - 2\mu$$

on the hypersurfaces $\{(q, p) | K_{\pm}(q, p) = 0\}$. The case of negative energy (that is, bounded orbits) is described by K_+ , the case of positive energy by K_- .

It is well known that, in three dimensions, this problem has SO(4) symmetry for negative energy and $SO_+(1,3)$ symmetry in the case of positive energy. We give explicit formulas for the actions of these groups on phase space $\mathbb{R}^3 \times \mathbb{R}^3$. Then we relate them to the regularizations of the Kepler problem by the Hopf map (Kustaanheimo–Stiefel [KS]) and by stereographic projection (Györgyi [Gy] /Moser [Mo] and Belbruno [Be]/Osipov [Os]).

To motivate the construction, we first consider the two dimensional situation. Here we identify phase space $\mathbb{R}^2 \times \mathbb{R}^2$ with $\mathbb{C} \times \mathbb{C}$. The symmetry groups SO(3) and $SO_+(1,2)$ have double covers SU(2) and SU(1,1), respectively. We show that the symmetries are related to the action of these groups on the momentum variable q by fractional linear transformations.

In three dimensions, we use the language of quaternions to describe the symmetries. In the case of negative energy, the symmetry group SO(4) has as double cover $SU(2) \times SU(2)$, which may be viewed as the group of pairs of quaternions of norm one. We identify \mathbb{R}^3 with the set of all "pure quaternions". The action of the group then is given by a formula similar to the two dimensional situation. In the case of negative energy, the symmetry group $SO_+(1,3)$ has as double cover $SL(2,\mathbb{C})$. We describe $SL(2,\mathbb{C})$ in terms of quaternions and again give the formula for the action on phase space.

Then we select a point on the energy hypersurface $\{(q, p) | K_{\pm}(q, p) = 0\}$. One gets a map from the symmetry group to the energy hypersurface by mapping a group element to the image of the chosen point under the action of the group element. We show that this map "is" the Kustaanheimo–Stiefel regularization. The regularizations of Györgyi/Moser and Belbruno/Osipov are based on the stereographic projections from the three dimensional sphere, respectively hyperboloid. By viewing the sphere as homogenous space for $SU(2) \times SU(2)$ and the hyperboloid as homogenous space for $SL(2, \mathbb{C})$, we get the relation between the Kustaanheimo–Stiefel regularization and the Györgyi/Moser resp. Belbruno/Osipov regularizations described by Kummer [Ku].

References:

[Be] J.Belbruno: Two Body motion under the inverse square central force and equivalent geodesic flows. Celestial Mechanics 15, 467-476 (1977)

[CB] R.Cushman, L.Bates: Global Aspects of Classical Integrable Systems. Birkhäuser Verlag 1997

[Gy] G.Györgi: Kepler's equation, Fock variables, Bacry's generators and Dirac brackets.Nuov. Cim. 53 A, 717-736 (1968)

[Ku] M.Kummer: On the regularization of the Kepler problem. Commun. Math. Phys. 84, 133-152 (1982)

[KS] P.Kustaanheimo, E.Stiefel: Perturbation theory of Kepler motion based on spinor regularization. J. Reine Angew. Math. **218**, 609-636 (1965)

[Mo] J.Moser: Regularization of Kepler's problem and the averaging method on a manifold.Comm. Pure Appl. Math. 23, 609-636 (1970)

[Os] Y.Osipov: The Kepler problem and geodesic flows in spaces of constant curvature. Celestial Mechanics 16, 191-208 (1977)