

An Intimate Gathering of Bosons



Peter van der Straten

Atom Optics (AOUD)

Honoursprogramma
Utrecht

September 29, 2009

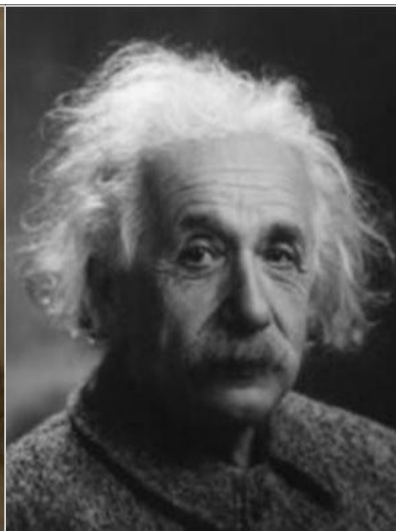


Universiteit **Utrecht**

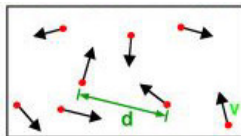
- 1 How to produce a Bose-Einstein condensate
 - Step 1: Light pressure
 - Step 2: Optical molasses
 - Step 3: Magnetic trapping
- 2 Observation of BEC
- 3 What to do with a Bose-Einstein condensate
 - Superradiant scattering
 - Collective excitations
 - Bosons *versus* Fermions
 - Superfluidity
 - Atom laser



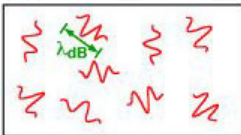
Bose-Einstein condensation



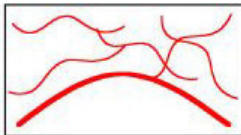
What is Bose-Einstein condensation (BEC)?



High Temperature T:
thermal velocity v
density d^{-3}
"Billiard balls"



Low Temperature T:
De Broglie wavelength
 $\lambda_{dB} = h/mv \propto T^{-1/2}$
"Wave packets"



$T = T_{crit}$:
Bose-Einstein
Condensation
 $\lambda_{dB} = d$
"Matter wave overlap"



$T = 0$:
Pure Bose
condensate
"Giant matter wave"



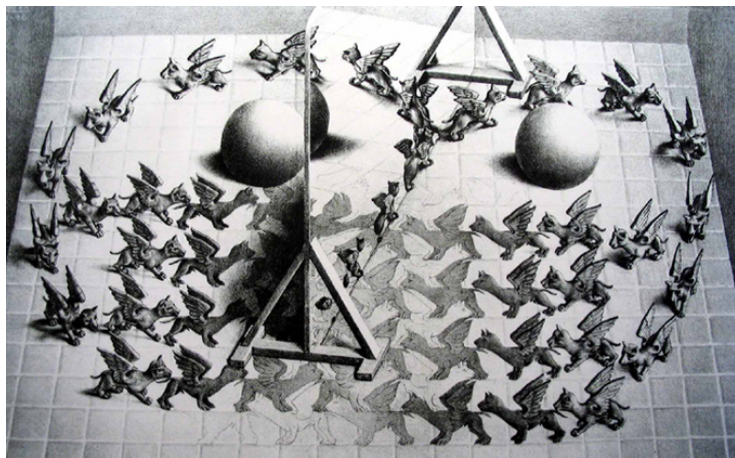
$$\rho = n\Lambda_{\text{deB}}^3 = 2.612375349 \dots$$

with

$$\Lambda_{\text{deB}} = \frac{h}{\sqrt{2\pi mk_B T}}$$



Bosons and Fermions



Usual examples: electrons, protons, neutron (fermions), photon (boson)

Example: ^3He ($I=1/2$) and ^4He ($I=0$), or ^6Li ($F=1/2, 3/2$) and ^7Li ($F=1$)



Pauli exclusion principle



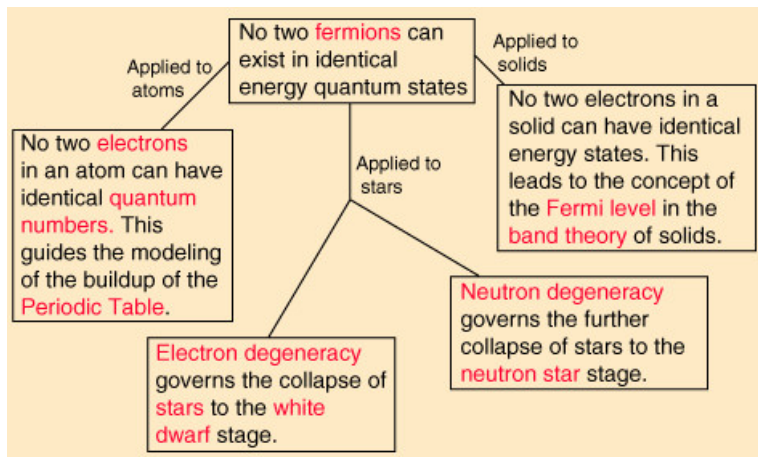
The Pauli exclusion principle is a quantum mechanical principle formulated by Wolfgang Pauli in 1925. It states that no two identical fermions may occupy the same quantum state simultaneously.

A more rigorous statement of this principle is that, for two identical fermions, the total wave function is anti-symmetric. For electrons in a single atom, it states that no two electrons can have the same four quantum numbers, that is, if n , l , and m_l are the same, m_s must be different such that the electrons have opposite spins.

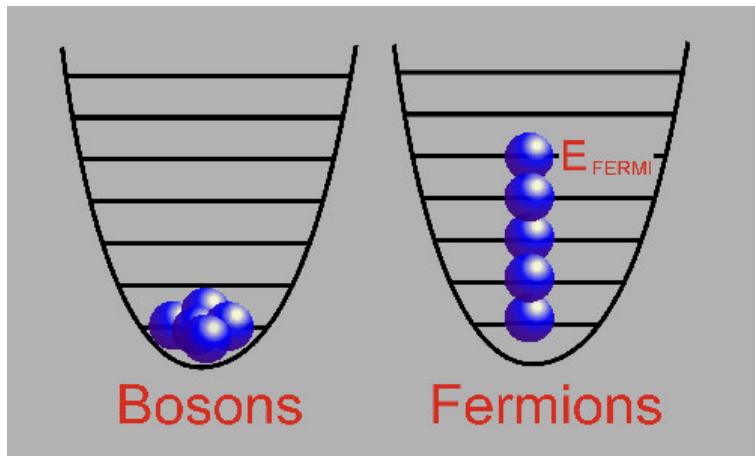
In relativistic quantum field theory, The Pauli principle follows from applying a rotation operator in imaginary time to particles of half-integer spin. It does not follow from any spin relation in nonrelativistic quantum mechanics.



Application of Pauli principle



Population of states

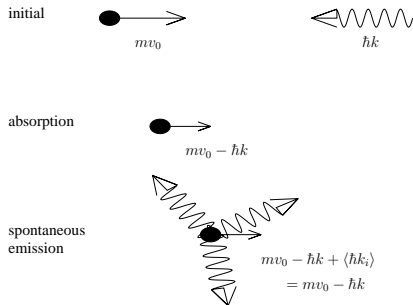


Outline

- 1 How to produce a Bose-Einstein condensate
 - Step 1: Light pressure
 - Step 2: Optical molasses
 - Step 3: Magnetic trapping
- 2 Observation of BEC
- 3 What to do with a Bose-Einstein condensate
 - Superradiant scattering
 - Collective excitations
 - Bosons *versus* Fermions
 - Superfluidity
 - Atom laser



Light pressure



recoil “kick”

$$v_r = \frac{\hbar k}{m} \approx 3 \text{ cm/s (Na)}$$

thermal

$$v \approx 1000 \text{ m/s}$$

$$N_{\text{stop}} \approx 33.000 \text{ fotons}$$

lifetime

$$\tau = 16 \text{ ns}$$

$$T_{\text{stop}} \approx 1 \text{ msec}$$

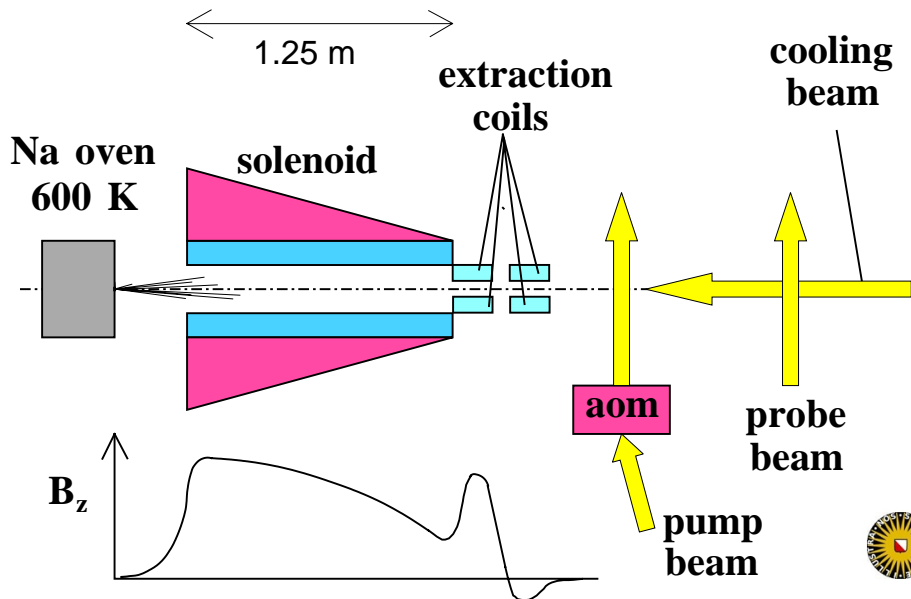
$$l_{\text{stop}} \approx 0.5 \text{ m}$$

acceleration

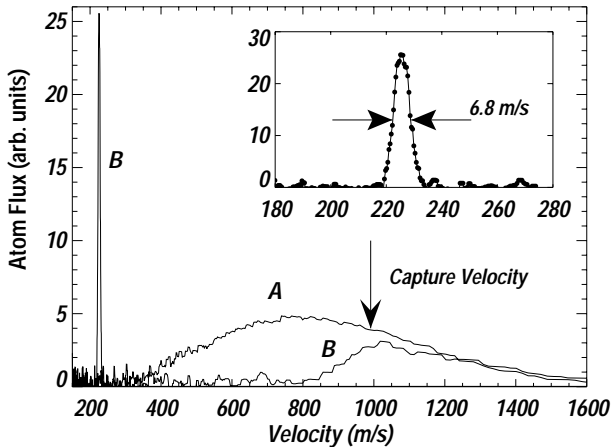
$$a \approx 9 \times 10^5 \text{ m/s}^2$$



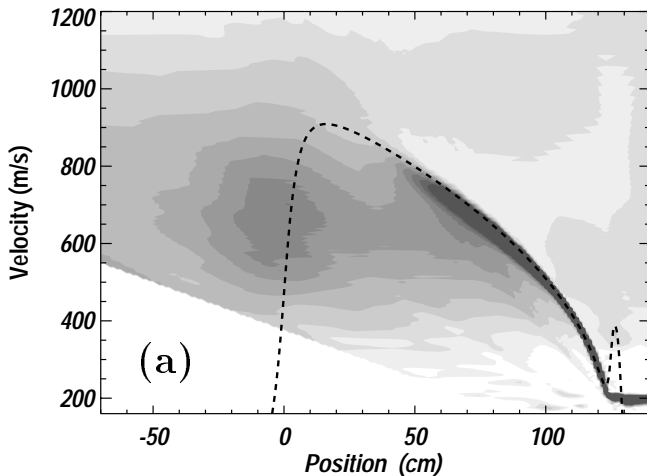
Zeeman technique



Velocity distribution



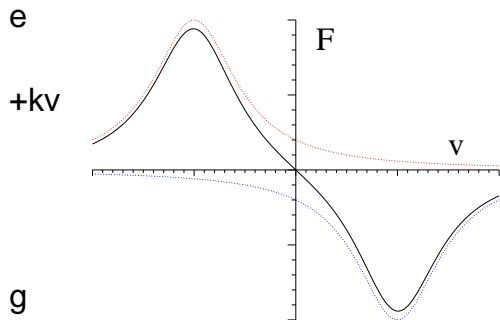
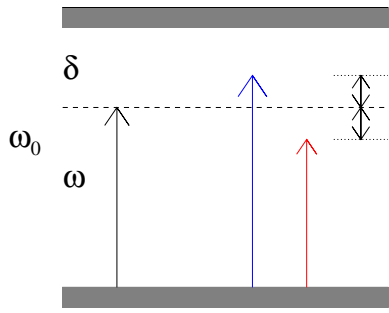
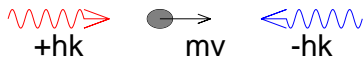
Funnel for atoms



Contour map of the velocity and position of atoms in the solenoid



Laser cooling

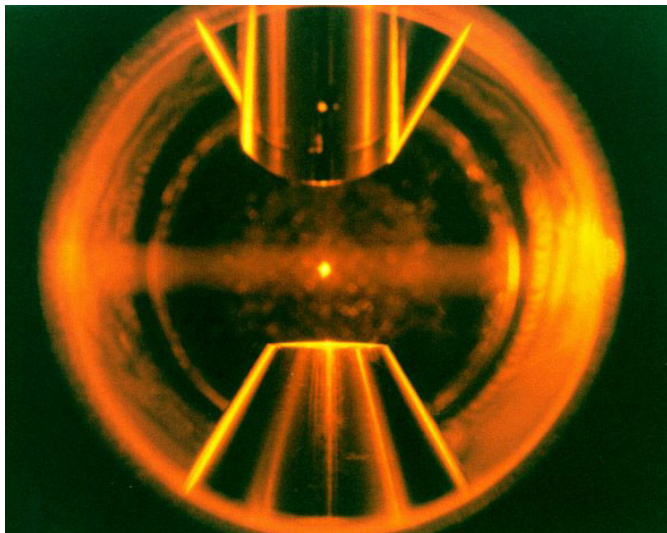


Cooling limit: Damping by Doppler tuning vs. heating by random recoil

$$kT_D = \frac{\hbar\Gamma}{2} \quad [\text{Na} : 240\mu\text{K}]$$



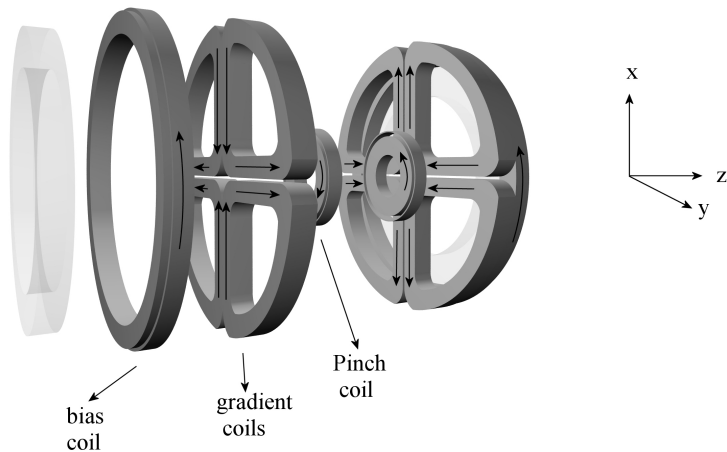
Cold Atoms



<D:/upload/Phys2000/bec/lascool4.html>



Magnetic Trap



Cloverleaf trap



Evaporation

- 1 Cools a cup of coffee
- 2 Cools apples by overtree sprinkling
- 3 Is used in technical water coolers
- 4 Globular clusters do it by evaporation of stars
- 5 Compound nuclei do it by evaporating neutrons
- 6 Atom coolers love it.

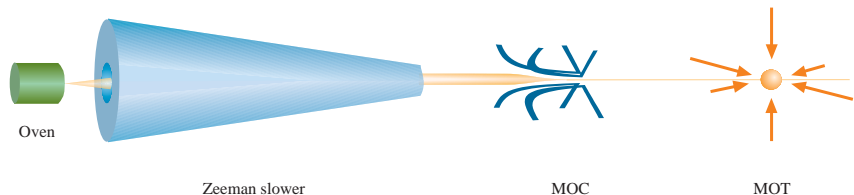
Evaporative cooling of a trapped gas is based on the preferential removal of atoms with an energy higher than the average energy, followed by thermalization of the gas by elastic collisions.

In order to force the cooling to proceed at a constant rate, the evaporation threshold may be lowered as the gas cools (forced evaporation).

D:/Upload/Phys2000/bec/xevap_cool.html



Experimental feasibility



Properties of the trapped atoms	N	n (atoms/cm ³)	T (μ K)
MOT	1.2×10^{10}	3×10^{11}	320
magnetic trap	8×10^9	3×10^{11}	340
evaporative cooling	1.5×10^8	8×10^{13}	0.3

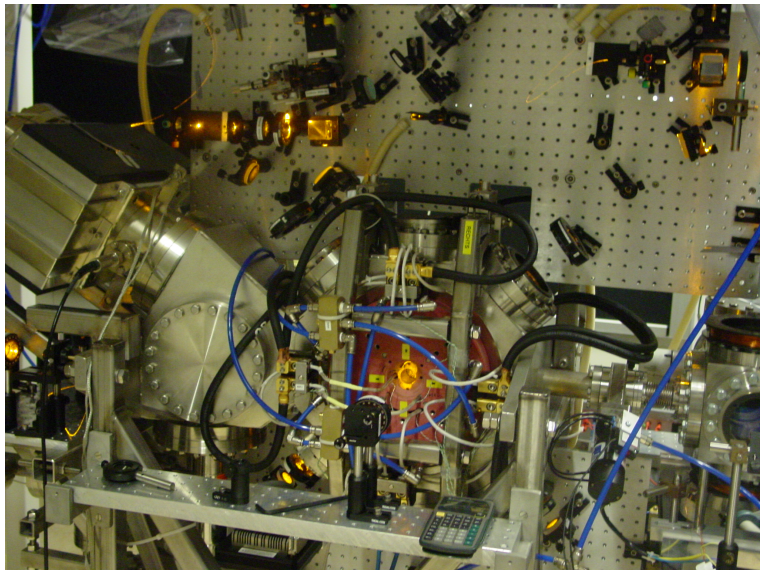


Outline

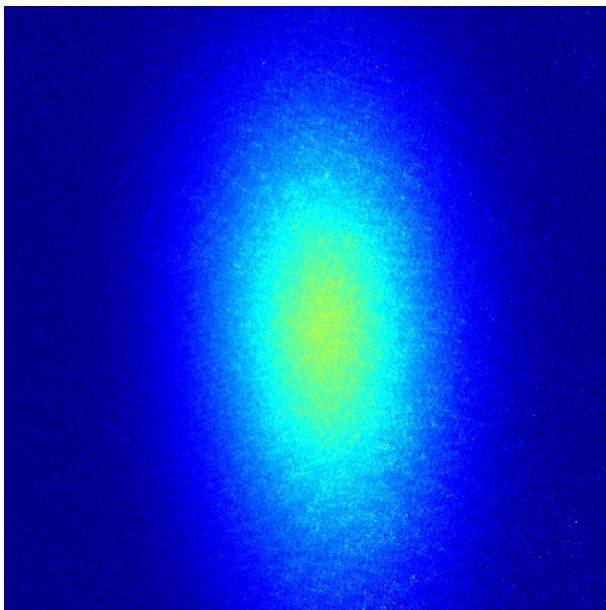
- 1 How to produce a Bose-Einstein condensate
 - Step 1: Light pressure
 - Step 2: Optical molasses
 - Step 3: Magnetic trapping
- 2 Observation of BEC
- 3 What to do with a Bose-Einstein condensate
 - Superradiant scattering
 - Collective excitations
 - Bosons *versus* Fermions
 - Superfluidity
 - Atom laser



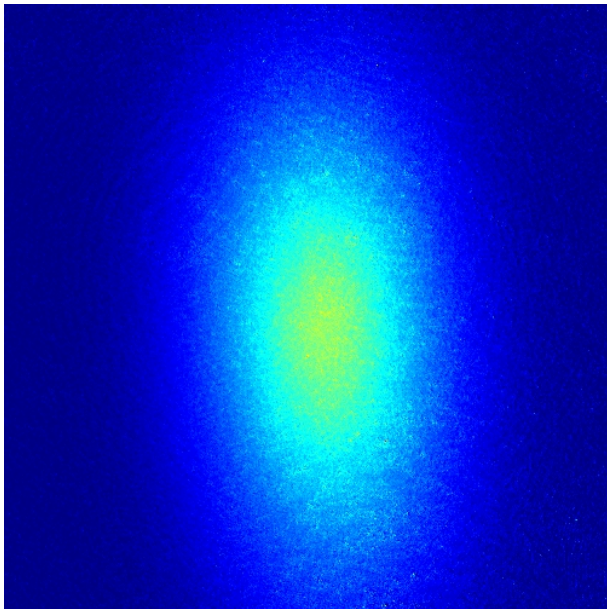
Current setup



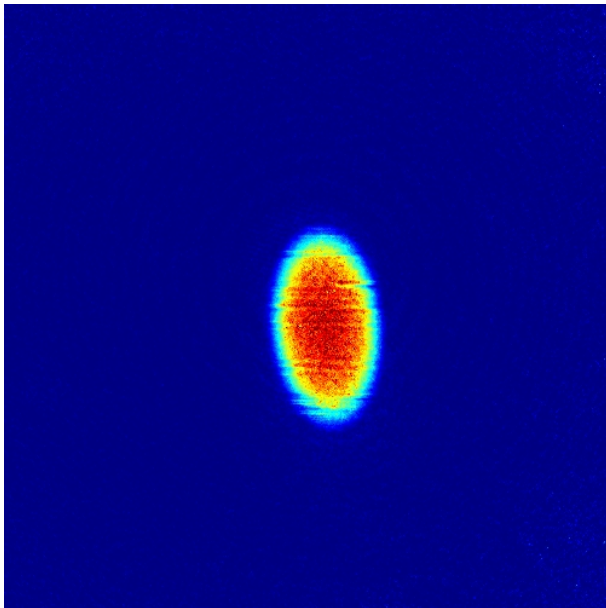
Bose-Einstein condensation—The last stages



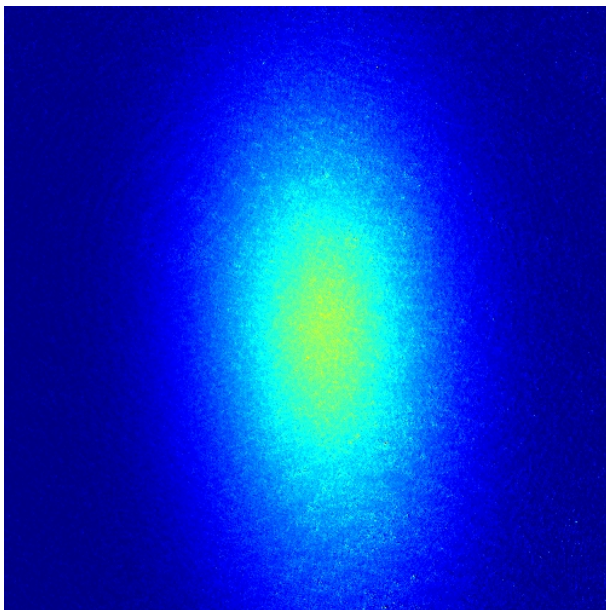
Bose-Einstein condensation—The last stages



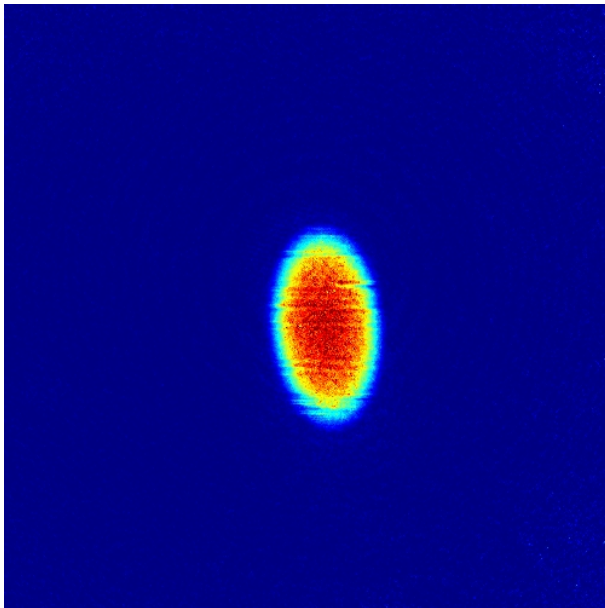
Bose-Einstein condensation—The last stages



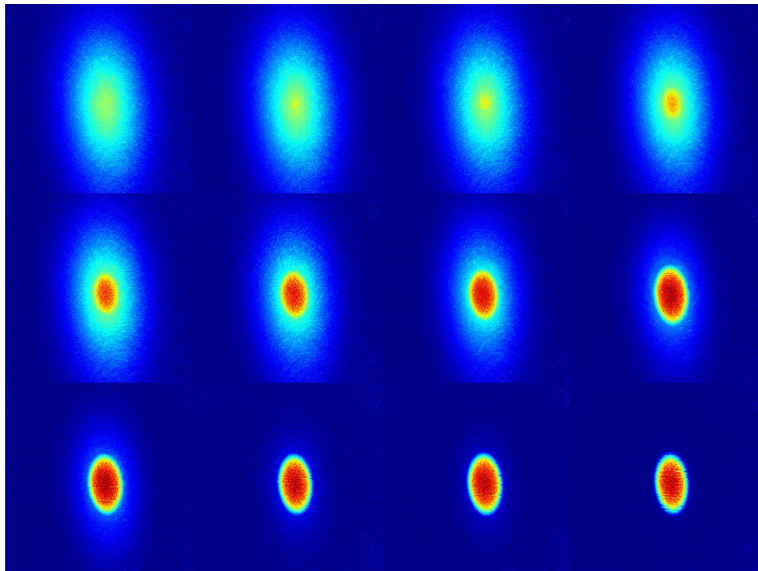
Bose-Einstein condensation—The last stages



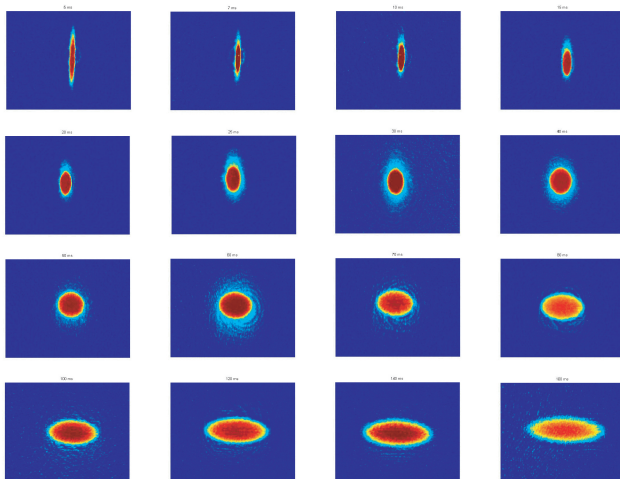
Bose-Einstein condensation—The last stages



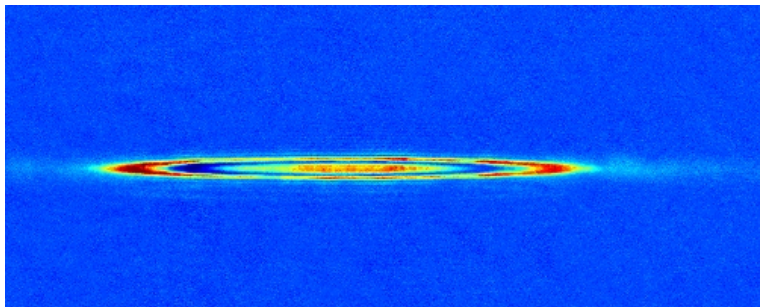
Bose-Einstein condensation—The last stages



Expansion of the cloud



Phase contrast imaging, “close” to resonance



$$\delta = -286 \text{ MHz}$$

Field-of-view $1.8 \times 0.7 \text{ mm}$

$$N = 1.5 \times 10^8 \text{ atoms}$$

Trap $96 \times 4 \text{ Hz}$

“in situ”



Characteristic values

Dark magneto-optical trap

Temperature	320 μK
Number of particles	1.2×10^{10} atoms
Density	3×10^{11} atoms/cm ³

Magnetic trap

Trap frequencies	$\nu_r=96$ Hz and $\nu_z=16 \rightarrow 1.08$ Hz
Number of particles	8×10^9 atoms
Elastic scattering rate	10 collisions/s

Bose-Einstein condensation

Evaporation ramp	40 s
Number of particles	2.5×10^8 atoms
Density	2.5×10^{14} atoms/cm ³
Chemical potential	3.5 kHz
Temperature	300 nK



Outline

- 1 How to produce a Bose-Einstein condensate
 - Step 1: Light pressure
 - Step 2: Optical molasses
 - Step 3: Magnetic trapping
- 2 Observation of BEC
- 3 What to do with a Bose-Einstein condensate
 - Superradiant scattering
 - Collective excitations
 - Bosons *versus* Fermions
 - Superfluidity
 - Atom laser



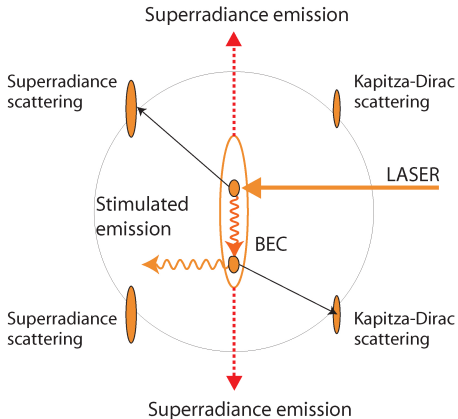
Topic One

Superradiant scattering

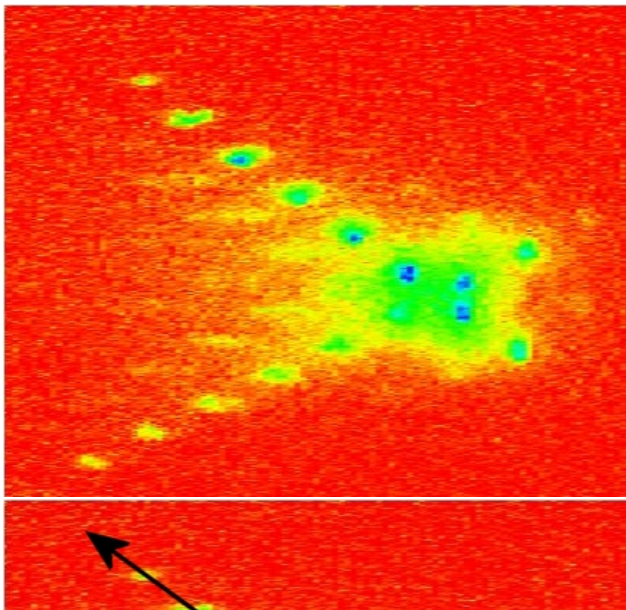
“Collective scattering of light from a Bose-Einstein condensate”



Superradiant backscattering–Setup



Superradiant scattering–Spectrum



Defining Entanglement

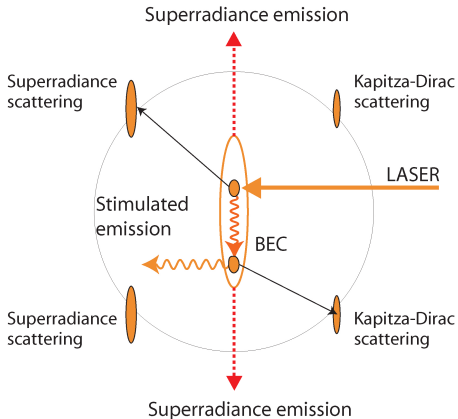
When two systems, of which we know the states by their respective representatives, enter into temporary physical interaction due to known forces between them, and when after a time of mutual influence the systems separate again, then they can no longer be described in the same way as before, viz. by endowing each of them with a representative of its own. I would not call that one but rather the characteristic trait of quantum mechanics, the one that enforces its entire departure from classical lines of thought. By the interaction the two representatives [the quantum states] have become entangled.



Erwin Schrödinger, 1935



Superradiant backscattering–Entanglement



Distinguishing Entanglement

VIEWPOINT

From Pedigree Cats to Fluffy-Bunnies

Jacob Dunningham, Alexander Rau, Keith Burnett*

We consider two distinct classes of quantum mechanical entanglement. The first “pedigree” class consists of delicate highly entangled states, which hold great potential for use in future quantum technologies. By focusing on Schrödinger cat states, we demonstrate not only the possibilities these states hold but also the difficulties they present. The second “fluffy-bunny” class is made up of robust states that arise naturally as a result of measurements and interactions between particles. This class of entanglement may be responsible for the classical-like world we see around us.

Dunningham *et al.*, *Science* **307**, 872 (2005)

are limited by more practical effects. Interferometry schemes, for example, usually use a stream of photons or atoms and are, therefore, normally limited by shot noise, where the measurement accuracy scales as $N^{-1/2}$. This conventional bound to measurement accuracy is a consequence both of the discrete



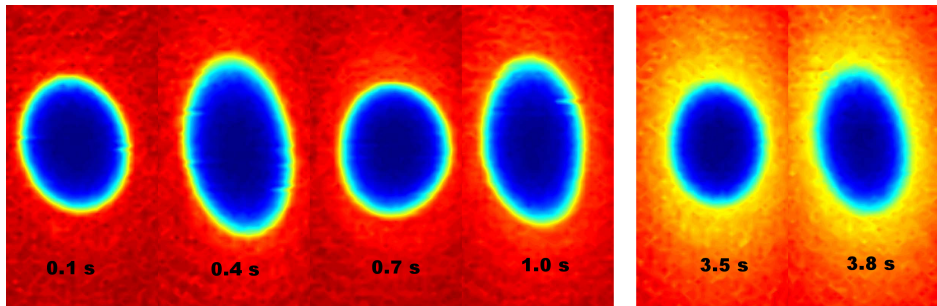
Topic Two

Collective excitations

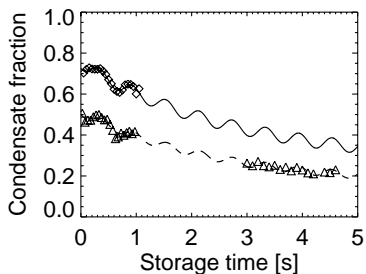
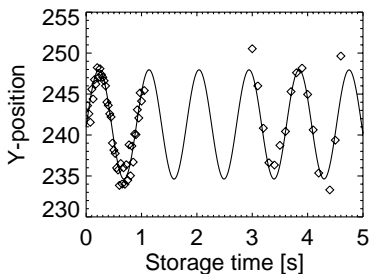
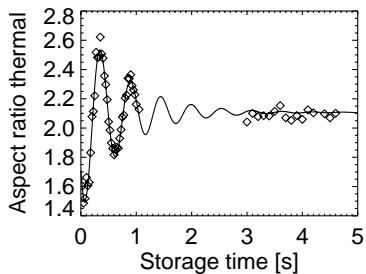
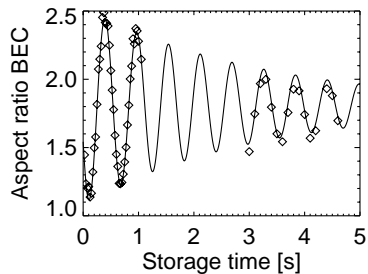
“Excitation of a mode of the condensate”

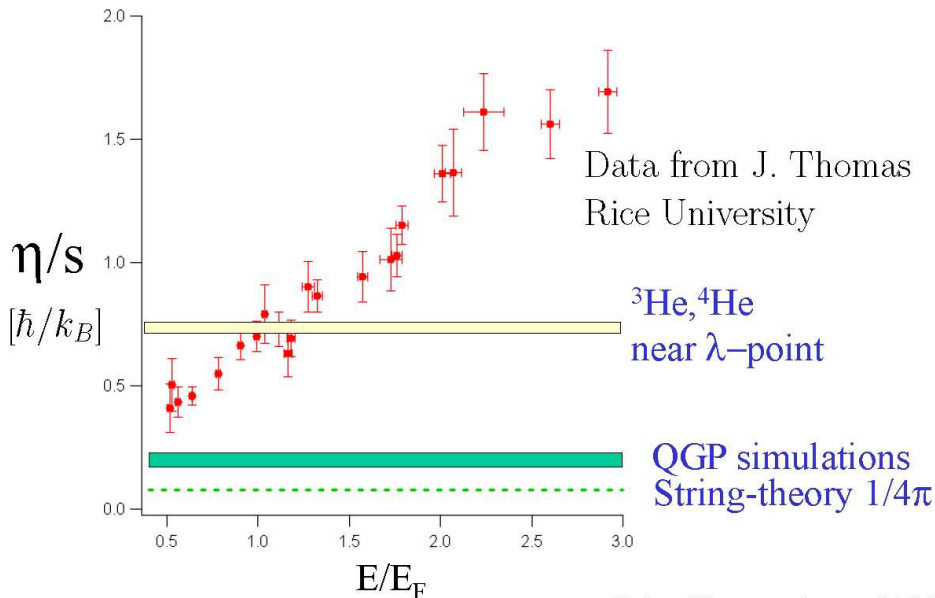


Quadrupole oscillations of a condensate



Damping of quadrupole oscillations



String theory, ^3He - ^4He , quark-gluon plasma and cold atoms

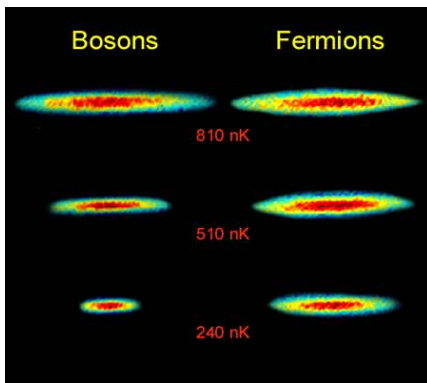
Topic Three

Bosons *versus* Fermions

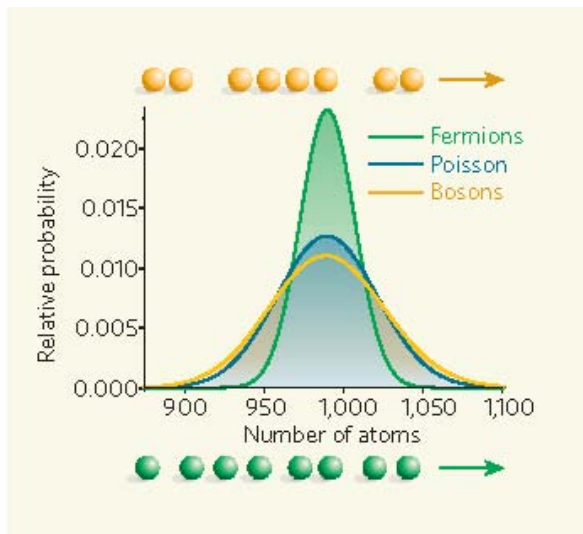
“How quantum statistics changes everything”



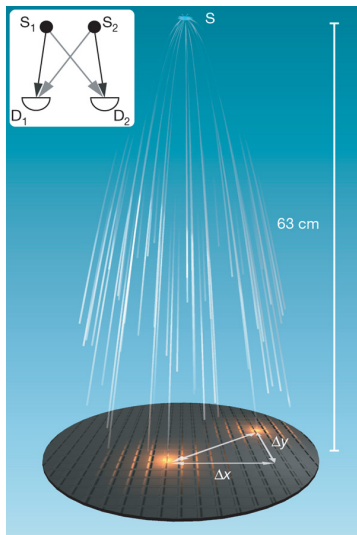
“Fermi” pressure



Bunching or anti-bunching



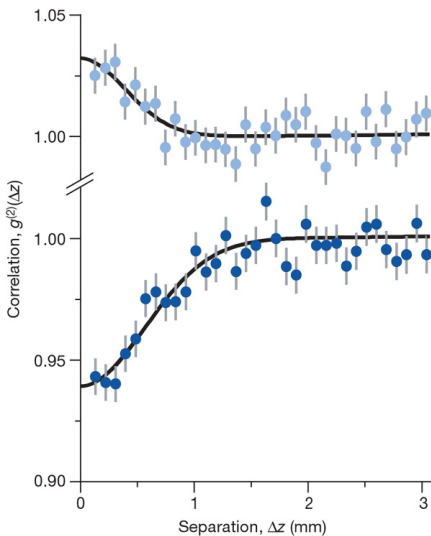
Bunching or anti-bunching



Jeltes *et al.*, Nature 2007



Bunching or anti-bunching



Jeltes *et al.*, Nature 2007



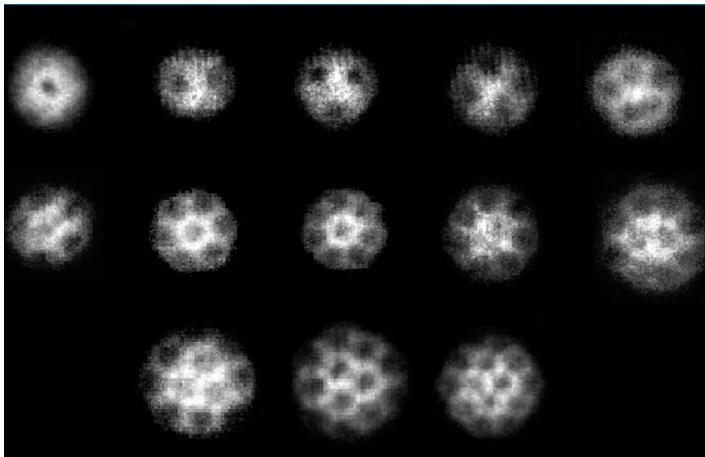
Topic Four

Prove of superfluidity

“Vortices in a quantum fluid”



Prove of superfluidity

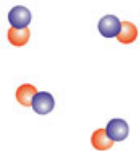


Madison *et al.*, Phys. Rev. Lett. **84**, 806 (2000)

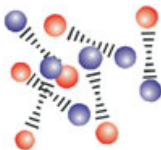


Cross-over between bosons and fermions

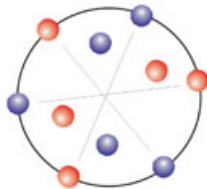
BEC ←————→ **BCS**



diatomic molecules



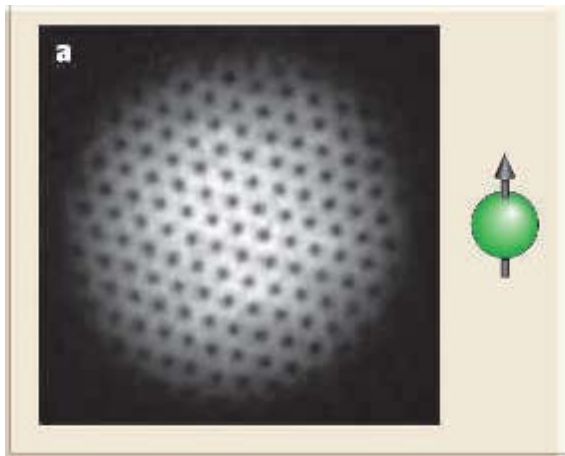
strongly interacting pairs



Cooper pairs



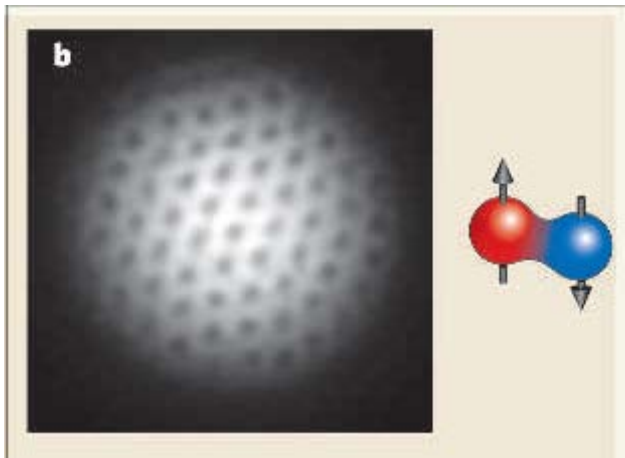
Prove of superfluidity for atoms



J. Abo-Shaer *et al.*, *Science* **292**, 476 (2001)



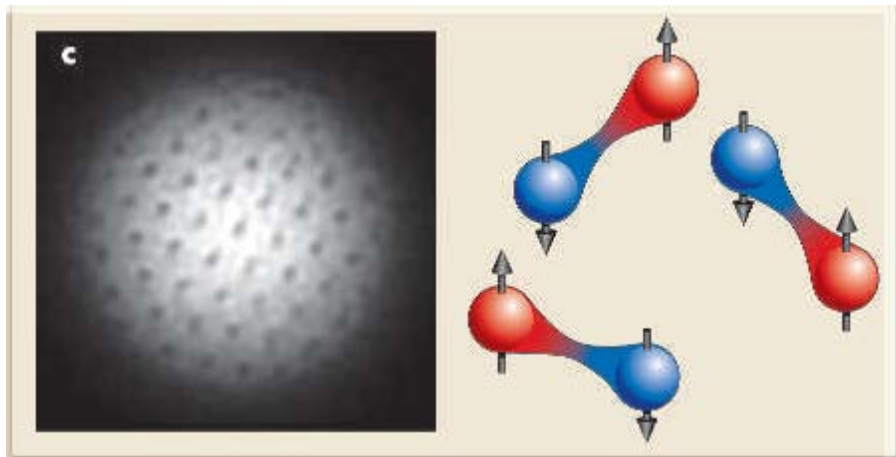
Prove of superfluidity for molecules



M. Zwierlein *et al.*, Nature **435**, 1047 (2005)



Prove of superfluidity for atom pairs



M. Zwierlein *et al.*, Nature **435**, 1047 (2005)



Topic Five

Atom laser

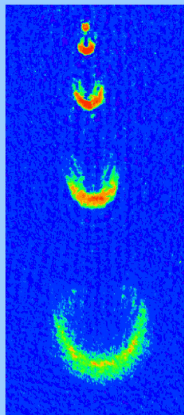
“The creation of a continuous flow of Bose-condensed atoms”



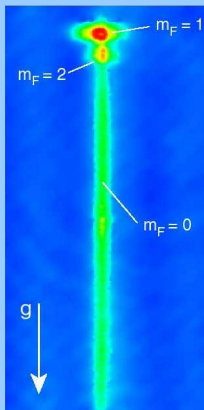
Atom lasers

Atom laser gallery

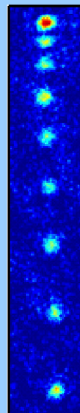
Height:
5, 2, 0.5, 1 mm



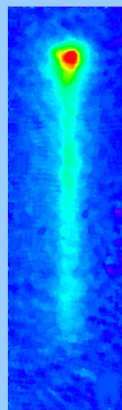
MIT '97



Munich '99



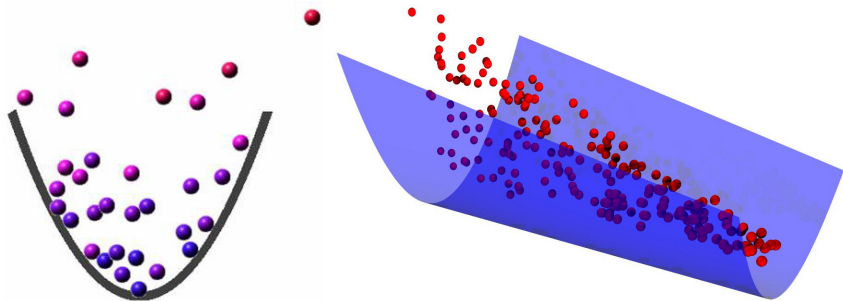
Yale '98



NIST '99



Evaporative cooling



Trap geometry vs. Beam geometry



“The rollercoaster”

